

Archival Lyman-Continuum and Theoretical Reionization Analysis vs Z

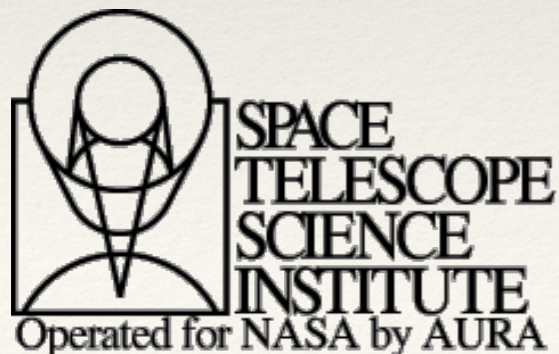
A.L.C.A.T.R.A.Z.

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Rogier Windhorst

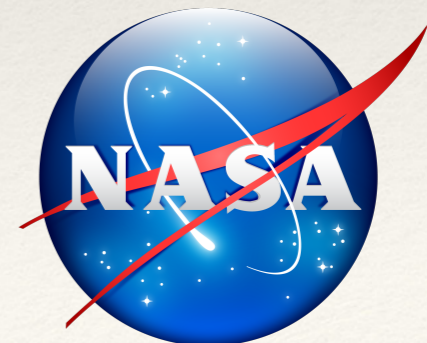
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Collaborators

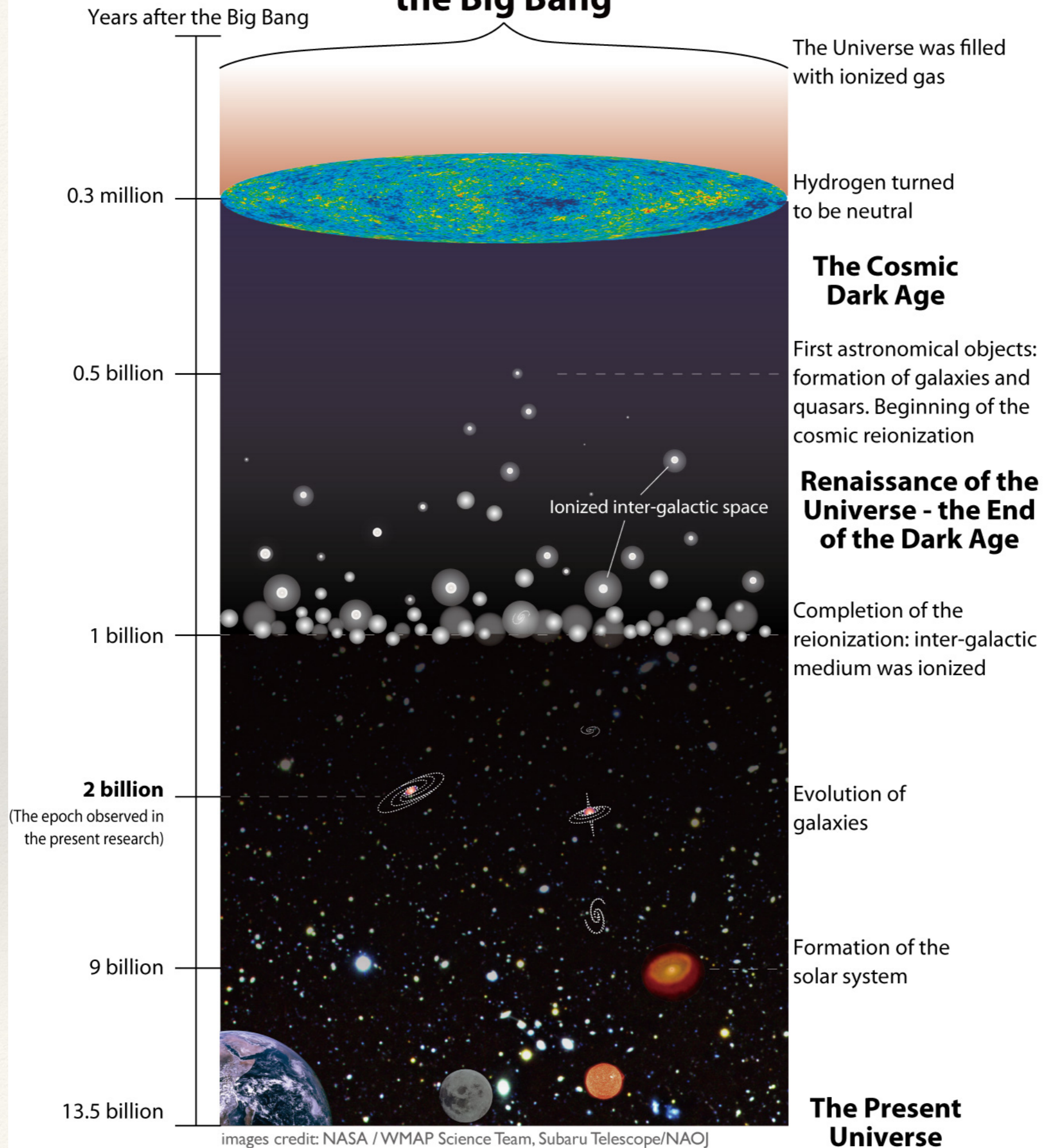
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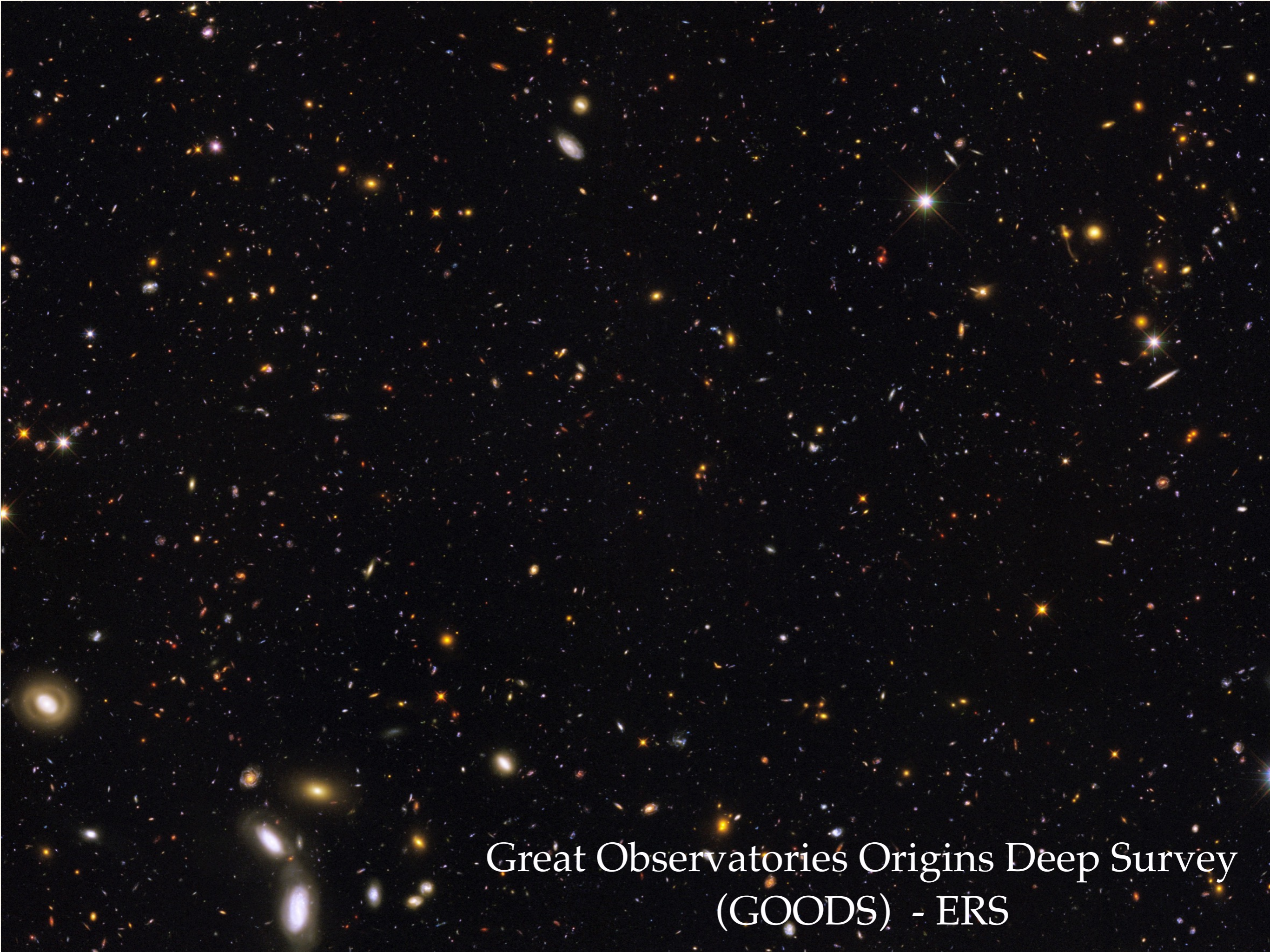


UiO : **University of Oslo**



the Big Bang





Great Observatories Origins Deep Survey
(GOODS) - ERS

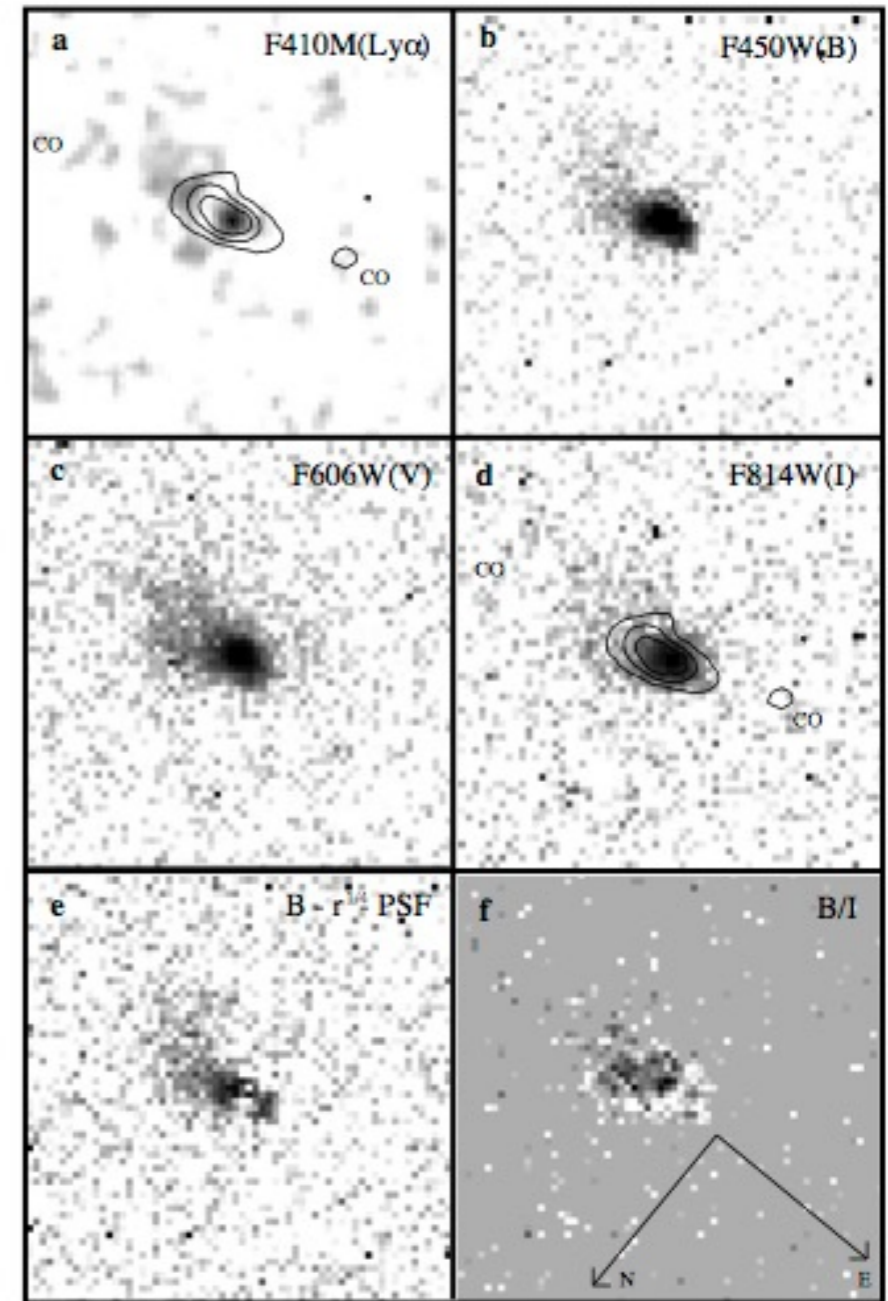
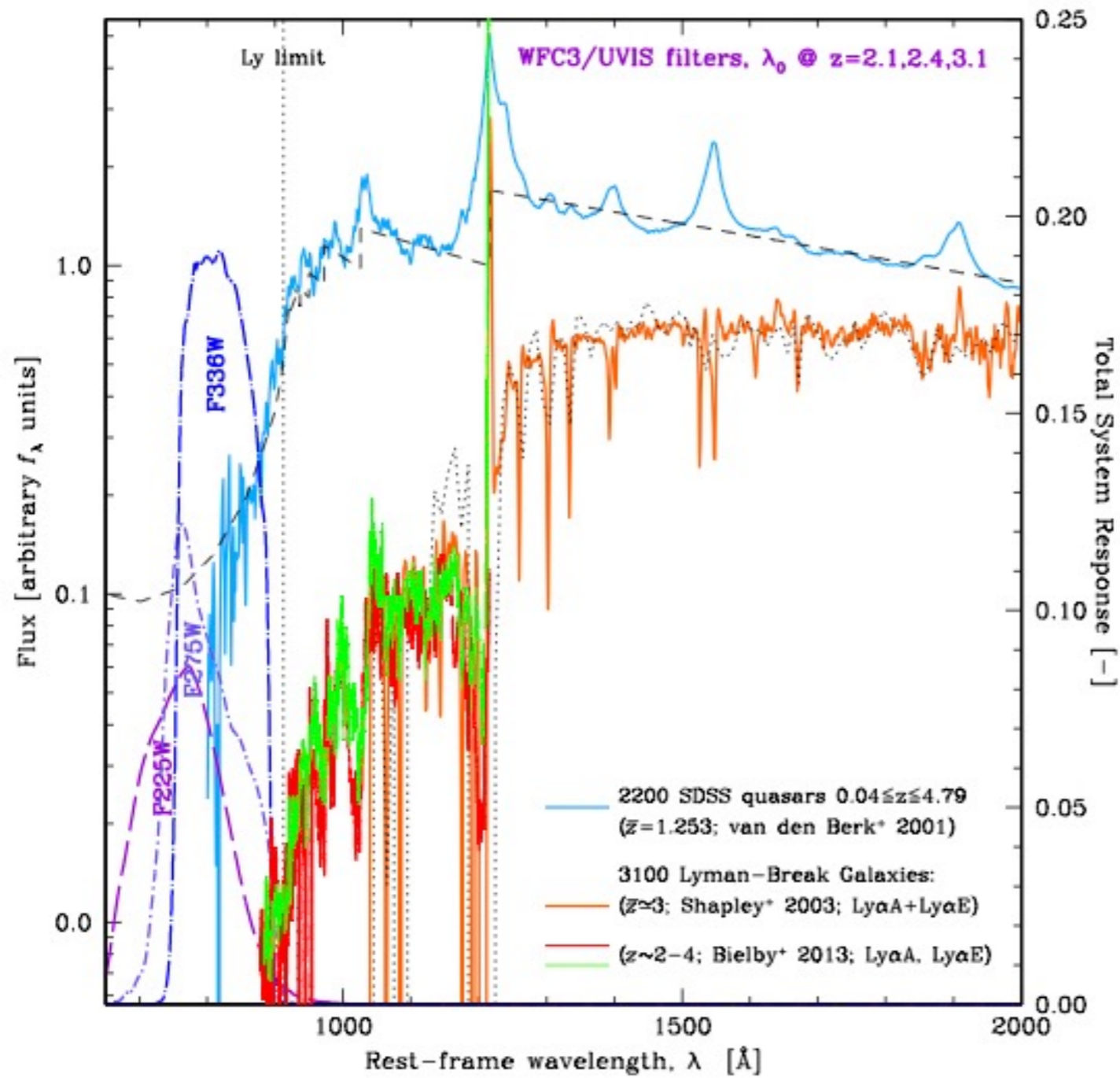


Fig. 1a [Right]: Composite rest-frame UV spectra from 2200 SDSS quasars (blue; van den Berk⁺ 2001) and 3100 LBGs [Shapley⁺ 2003 (orange); Bielby⁺2013 (red/green)]. WFC3 UV throughputs in F225W, F275W, F336W sample LyC for objects at $z=2.26, 2.45, 3.07$. **Fig. 1b [RIGHT]:** Radio galaxy with weak AGN at $z=2.390$ (Windhorst⁺ 1998) in WFPC2 F410M (redshifted Ly α), B, V and I (longwards of Ly α). Bottom panels show the PSF-subtracted B-band image and (B-I) color gradient. In all panels, a blue cloud is visible towards the upper-left, along which LyC may escape. This is aligned with its radio source and CO- clouds.

Gold vs Silver

TABLE 2
LYC STACK SUMMARIES OF GOLD SAMPLE (HIGH QUALITY SPECTRA)

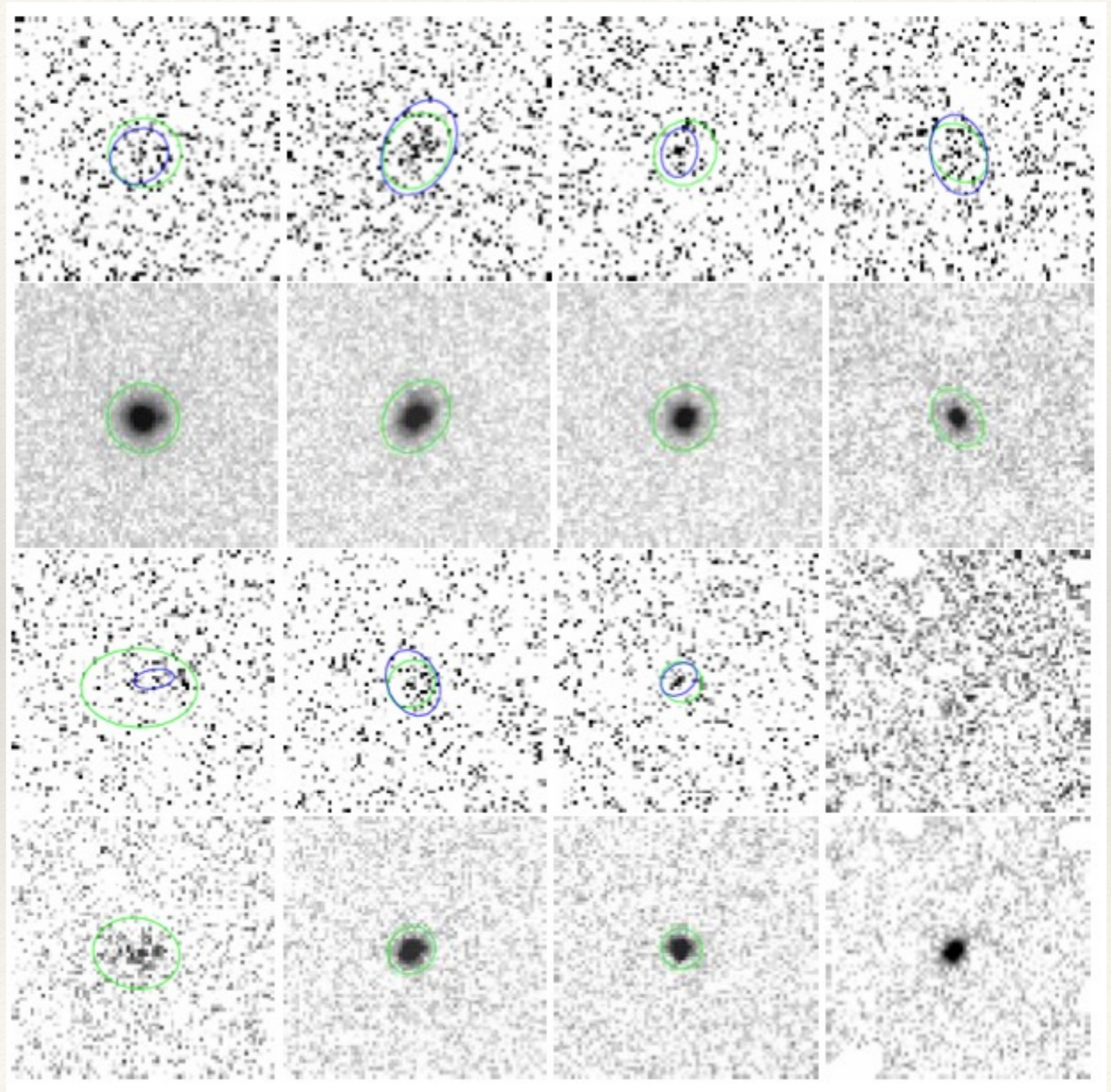
Filter	z^a	N^b	$m(\text{LyC,obs})^c$	SNR^d	$m(\text{LyC},1500\text{\AA})^e$	$\text{SNR}(\text{LyC},1500\text{\AA})^f$	$m(\text{UV}_{1500\text{\AA}})^g$	$\text{SNR}(\text{UV}_{1500\text{\AA}})^h$	$\text{LyC}(\text{Aper}_{area})^i$	$\text{UV}_{1500\text{\AA}}(\text{Aper}_{area})^i$	f_{esc}^k
Gold Sample: All Objects											
F225W	2.291 – 2.449/2.369	14	> 30.76	0.08 σ	30.76	0.08 σ	24.49	195 σ	2.132	2.132	
F275W	2.450 – 3.009/2.653	18	30.61	7.50 σ	29.95	0.17 σ	24.86	159 σ	1.687	2.145	
F336W	3.076 – 3.917/3.451	16	30.02	13.45 σ	30.16	0.25 σ	24.57	123 σ	0.679	1.522	
F435W	4.800 – 5.977/5.230	6	> 99.0	0.28 σ	99.0	0.28 σ	25.72	25 σ	1.188	1.188	
Gold Sample: AGN only											
F225W	2.291 – 2.291/2.291	1	> 34.68	0.05 σ	34.68	0.05 σ	25.93	10.67 σ	6.260	6.260	
F275W	2.470 – 3.009/2.697	7	31.01	6.66 σ	29.70	0.17 σ	24.98	124 σ	1.740	1.063	
F336W	3.217 – 3.474/3.346	2	29.14	9.32 σ	29.72	0.22 σ	23.99	125 σ	0.518	0.773	
Gold Sample: Galaxies Only											
F225W	2.291 – 2.449/2.369	13	> 30.43	0.08 σ	30.43	0.08 σ	24.44	201 σ	2.083	2.098	
F275W	2.450 – 2.975/2.625	11	30.64	8.17 σ	30.59	0.06 σ	24.80	119 σ	0.484	2.675	
F336W	3.076 – 3.917/3.466	14	29.98	12.57 σ	30.32	0.16 σ	24.69	97 σ	0.662	1.668	
LyC Stack Summaries of Silver Sample (High and Intermediate Quality Spectra)											
Silver Sample: All Objects											
F225W	2.262 – 2.449/2.360	32	31.57	9.20 σ	30.78	0.08 σ	24.59	238 σ	1.588	2.311	
F275W	2.450 – 3.009/2.650	33	29.71	14.17 σ	30.03	0.21 σ	24.97	184 σ	3.120	2.243	
F336W	3.076 – 4.149/3.511	30	30.70	11.48 σ	30.77	0.22 σ	24.58	149 σ	0.880	1.929	
F435W	4.379 – 5.979/5.081	13	27.50	4.69 σ	27.71	3.29 σ	25.39	45 σ	1.923	1.188	
Silver Sample: Galaxies Only											
F225W	2.262 – 2.449/2.359	30	31.53	9.41 σ	30.76	0.08 σ	24.55	201 σ	1.460	2.345	
F275W	2.450 – 2.975/2.638	26	29.90	9.77 σ	30.15	0.13 σ	24.98	154 σ	2.989	2.530	
F336W	3.076 – 4.149/3.522	28	30.79	10.30 σ	30.58	0.25 σ	24.65	131 σ	0.784	2.000	

^a Redshift range of galaxies included in stack/average redshift of stack.
^b Number of galaxies with high or intermediate quality spectroscopic redshifts used in stack.
^c Observed total magnitude of LyC of stack ($AB_{mag}^{m^2}$)
^d Signal to noise ratio of LyC detection in stack
^e Observed total magnitude of LyC of stack (AB_{mag}) measured in $\text{UV}_{1500\text{\AA}}$ stack aperture (green; see Fig. 5–8)
^f Signal to noise ratio of LyC detection in stack measured in $\text{UV}_{1500\text{\AA}}$ stack aperture
^g Observed rest-frame $\text{UV}_{1500\text{\AA}}$ magnitude of stack
^h Signal to noise ratio of LyC detection in stack measured in $\text{UV}_{1500\text{\AA}}$ stack aperture (green; see Fig. 5–8)
ⁱ Size of LyC aperture in arcseconds
^j Size of $\text{UV}_{1500\text{\AA}}$ aperture in arcseconds
^k Escape fraction of LyC where $f_{esc} = \frac{(f_{1500}/f_{LyC})_{stel}}{(f_{1500}/f_{LyC})_{obs}}$

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From Left to Right: Image stacks of Gold +Silver sample in the HST WFC3 filters F225W at $z=2.26-2.47$ (left), F275W at $z=2.47-3.08$ (middle left), F336W at $z=3.08-4.35$ (middle right), and ACS F435W filter at $z=4.35-6$ (right columns): (Row 1): LyC (Row 2): UVC for All Gold+Silver objects; (Row 3): LyC & (Row 4): UVC for (Gold) AGN. Blue ellipses indicate SExtractor LyC apertures using a 1σ detection criterion. Green ellipses are SExtractor MAG AUTO elliptical apertures, using the centroids and aperture of the restframe 1500\AA UV-continuum (UVC) stacks. All image-stacks are $6.''4 \times 6.''4$.

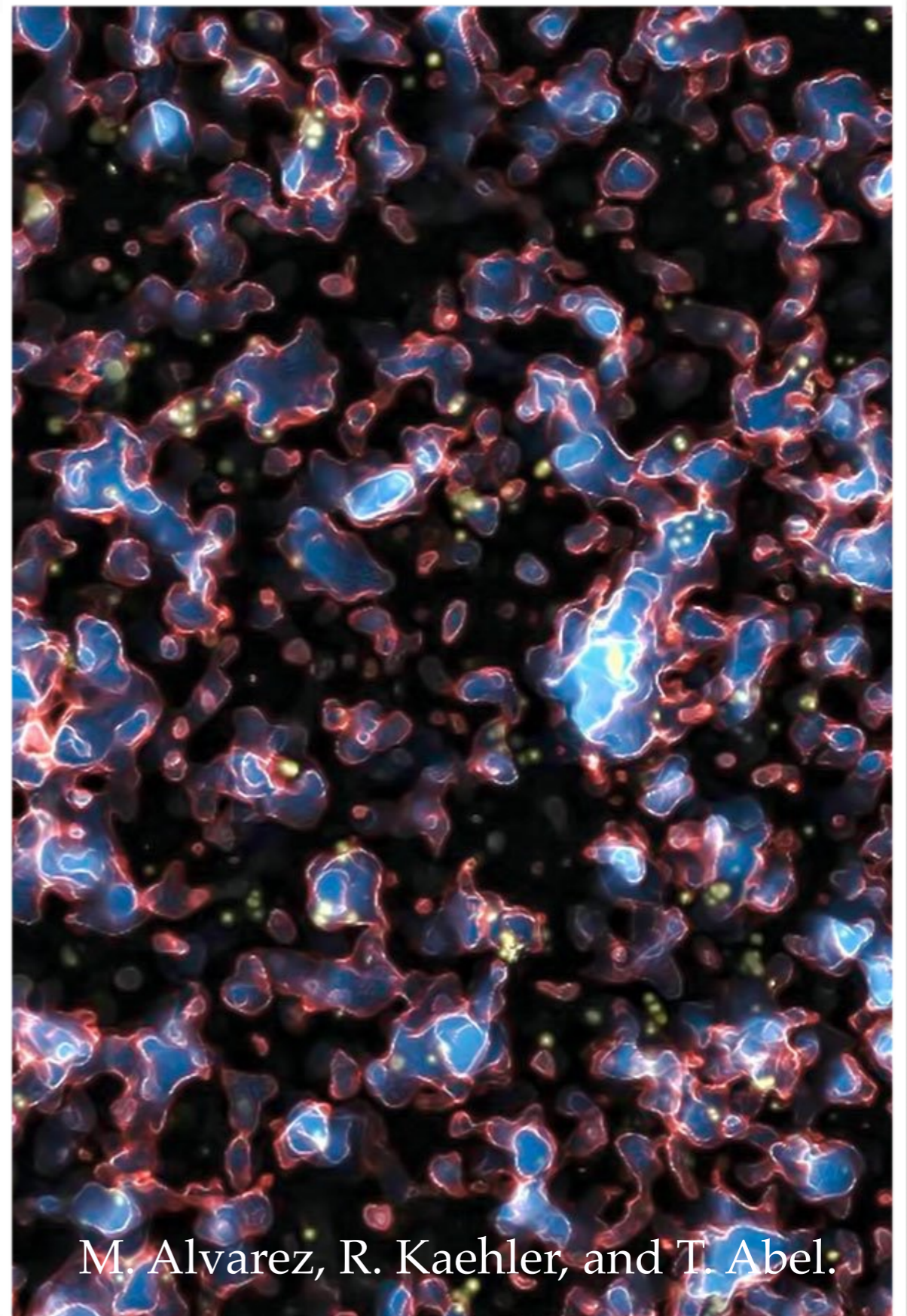
The remarkable result is that, despite very rigorous spectroscopic sample vetting, significant LyC is seen at most redshifts and for all samples.



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A.L.C.T.R.A.Z. 2015-16

- Proposal resubmitted for Hubble Cycle 23 and NSF funding, pending approval.
- Further analysis of HUDF to check results from ERS
- Study galaxy assembly with JWST
- Study Super-Massive Black-Hole Growth with JWST
- Study First Light and Reionization with JWST
- Study best groups and clusters lenses to find $z > 10$



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